

The structure of ^{12}N using $^{11}\text{C} + \text{p}$ resonance scattering

K. Peräjärvi¹, D. Lee^{1,2}, V.Z. Goldberg³, F.Q. Guo^{1,4}, D.M. Moltz⁴, J.P. O'Neil⁵, J. Powell¹, and Joseph Cerny^{1,4}

¹ Nuclear Science Division, Lawrence Berkeley National Laboratory, Berkeley, CA 94720, USA

² Department of Nuclear Engineering, University of California, Berkeley, CA 94720, USA

³ Texas A & M University, Cyclotron Institute, College Station, Texas 77843, USA

⁴ Department of Chemistry, University of California, Berkeley, CA 94720, USA

⁵ Life Sciences Division, Lawrence Berkeley National Laboratory, Berkeley, CA 94720, USA

The level structure of the drip-line nucleus ^{12}N has been studied in the past using conventional beams, see [1] and references therein. More recent references [2, 3] are basically extensions of those earlier works. Interest in the nuclear structure of ^{12}N is mainly related to the idea that many low-lying levels in ^{12}N should manifest one-particle-one-hole configurations, and therefore their features provide a test (and parameters) for shell model calculations.

In addition to the nuclear physics interest, studies involving ^{12}N around its $^{11}\text{C} + \text{p}$ threshold at 0.601 MeV are often also driven by nuclear astrophysics interests [4-8]. Namely, to be able to accurately determine the astrophysical rate of the $^{11}\text{C}(\text{p},\gamma)^{12}\text{N}$ reaction, detailed knowledge of the low-lying level structure of ^{12}N is also required. The $^{11}\text{C}(\text{p},\gamma)^{12}\text{N}$ reaction is associated with hot pp chains that might be able to bypass the triple alpha process in producing CNO material in low metallicity stars [9]. The ^{12}N excitation region in the vicinity of the $^8\text{B} + ^4\text{He}$ threshold at 8.008 MeV is also important for astrophysics due to the formation of ^{11}C in the $^8\text{B}({}^4\text{He},\text{p})^{11}\text{C}$ reaction [9]. Favorable states in ^{12}N close to this threshold could strongly enhance the corresponding reaction rate. The $^8\text{B}({}^4\text{He},\text{p})^{11}\text{C}$ reaction has been experimentally studied in [10]. That measurement was done using the inverse kinematics reaction, i.e., it utilized a radioactive ^{11}C beam. Resonant states between 8.7 and 9.9 MeV in ^{12}N were probed, and no resonant structures were reported.

As a result of these past studies, we can conclude that the level structure in ^{12}N is relatively well established up to the first 3^- state at about 3.1 MeV. However, above that energy, the level scheme of ^{12}N is not yet complete. The present work extends our knowledge of the level structure of ^{12}N by covering the excitation energy interval from 2.2 to 11.0 MeV using the $^{11}\text{C} + \text{p}$ resonance interaction with thick targets and inverse kinematics.

Ninety and 125 MeV ^{11}C beams were obtained using the Berkeley Experiments with Accelerated Radioactive Species (BEARS) coupled cyclotron system [11], which can produce ^{11}C beams at 2×10^7 ions/s on target, an about 100 times stronger ^{11}C beam than, for example, was available for [10]. As an example of the quality of the experimental data obtained in the present work, Fig. 1 presents the zero degree $^{11}\text{C} + \text{p}$ excitation function. To convert the E_{cm} energy scale in Fig. 1 to excitation energy, 0.601 MeV should be added to it. The R-matrix analysis [12] of the data is currently in progress.

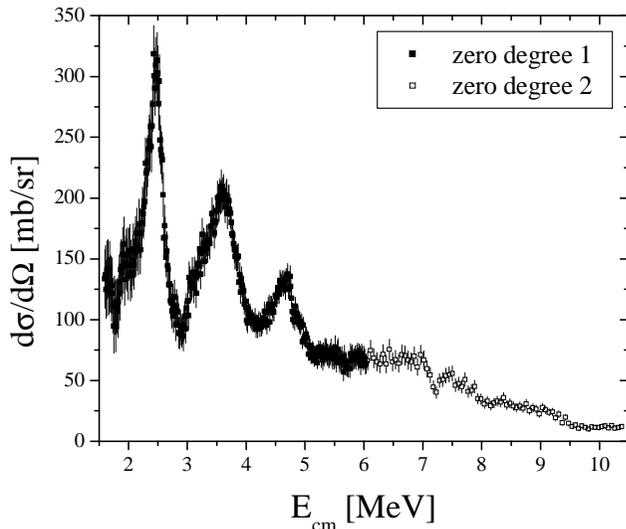


Fig. 1. The excitation function for $^{11}\text{C} + \text{p}$ elastic resonance scattering from 2.2 to 11.0 MeV in excitation energy, measured at zero degrees. Two different setups were used. Setup 1 covered 1.6 to 6.0 MeV and setup 2 covered 5.9 to 10.4 MeV in the c.m. system. See text.

REFERENCES

- [1] F. Ajzenberg-Selove, Nucl. Phys. A 506 (1990) 1.
- [2] M.N. Harakeh et al., Nucl. Phys. A 577 (1994) 57c.
- [3] B.D. Anderson et al., Phys. Rev. C 54 (1996) 237.
- [4] A. Lefebvre et al., Nucl. Phys. A 592 (1995) 69.
- [5] A. Galindo-Uribarri et al., Frontiers of Nuclear Structure, American Institute of Physics, vol. CP656 (2003) 323.
- [6] T. Teranishi et al., Phys. Lett. B 556 (2003) 27.
- [7] X. Tang et al., Phys. Rev. C 67 (2003) 015804.
- [8] W. Liu et al., Nucl. Phys. A 728 (2003) 275.
- [9] M. Wiescher et al., Astrophys. J. 343 (1989) 352.
- [10] K.E. Rehm et al., Nucl. Phys. A 746 (2004) 354c.
- [11] J. Powell et al., Nucl. Instrum. Methods Phys. Res. A 455 (2000) 452.
- [12] A.M. Lane and R.G. Thomas, Rev. Mod. Phys. 30 (1958) 257.